# Specifactory jancami andrewmarkwick-kemper

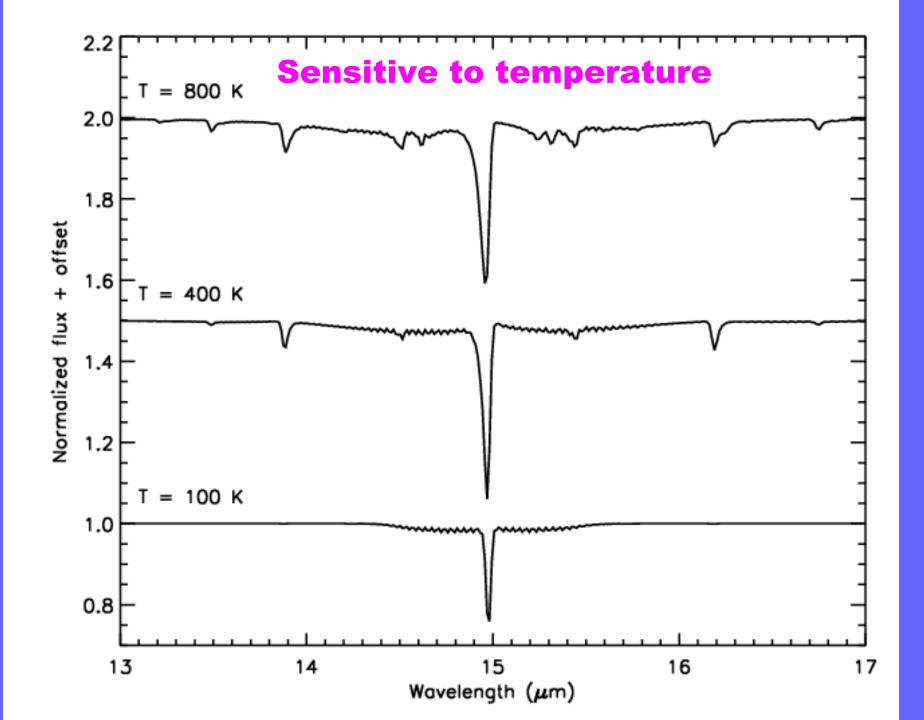
A dynamic database of molecular model spectra

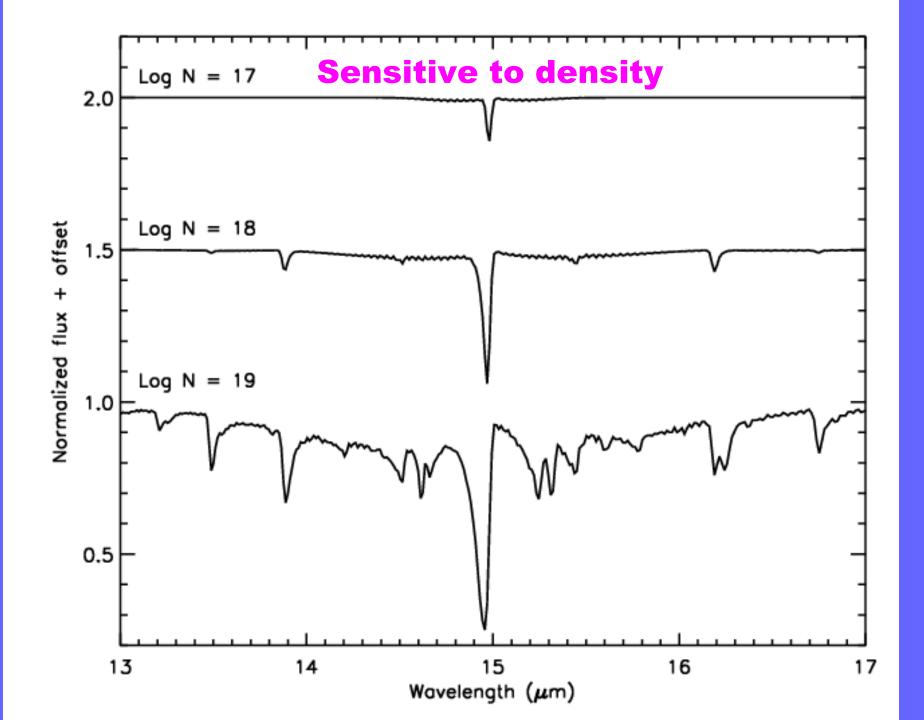
# **Outline**

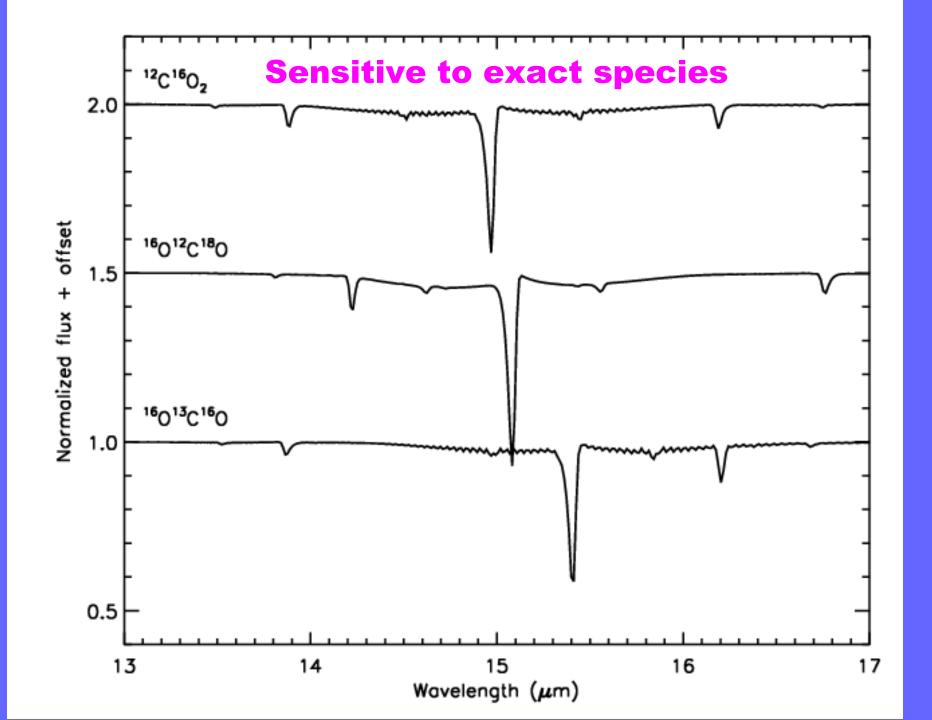
- Motivation / introduction
- The model calculations
- The web interface
- Current status
- Some examples
- Future prospects & needs

## **Motivation**

- We live in a molecular universe I molecules detected in/around young stars, old stars, interstellar medium, planetary atmospheres, comets, entire galaxies, ...
- Molecules are observed from UV to sub-mm wavelengths
- Important diagnostic tools & probes !!







# **Extracting physical parameters**

#### Identification of species

- Where to find frequencies & line strengths?
- Which one of those millions of molecules?
- What band profile?

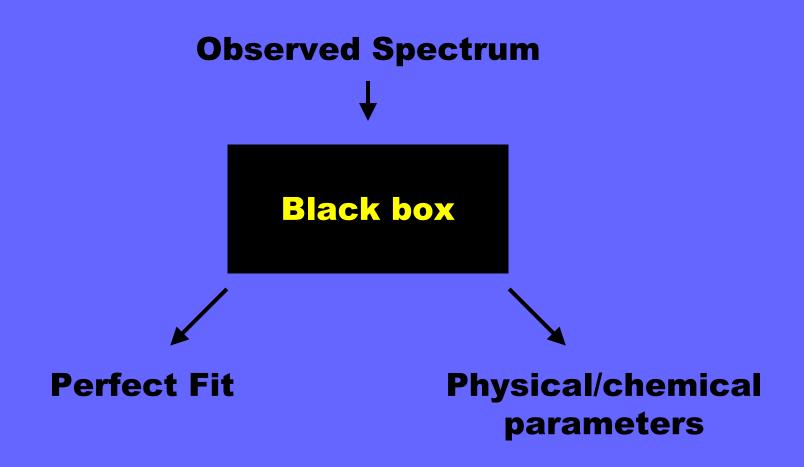
#### Spectral dependence on physical parameters

- Some knowledge of molecular spectroscopy
- Some knowledge of radiative transfer

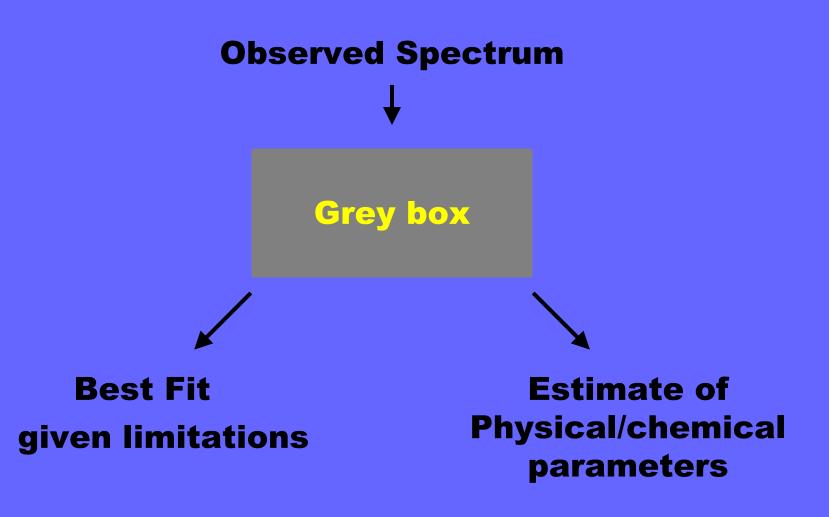
#### **Model fit**

- All of the above, plus some statistics

## In an ideal world...



## **Our** aim



**Black box** 

**Grey box** 

All possible molecular species

What's available

All possible transitions

What's available

**Full non-LTE radiative transfer** 

sotherm, LTE

Various geometries

Slab

**All possible instruments** 

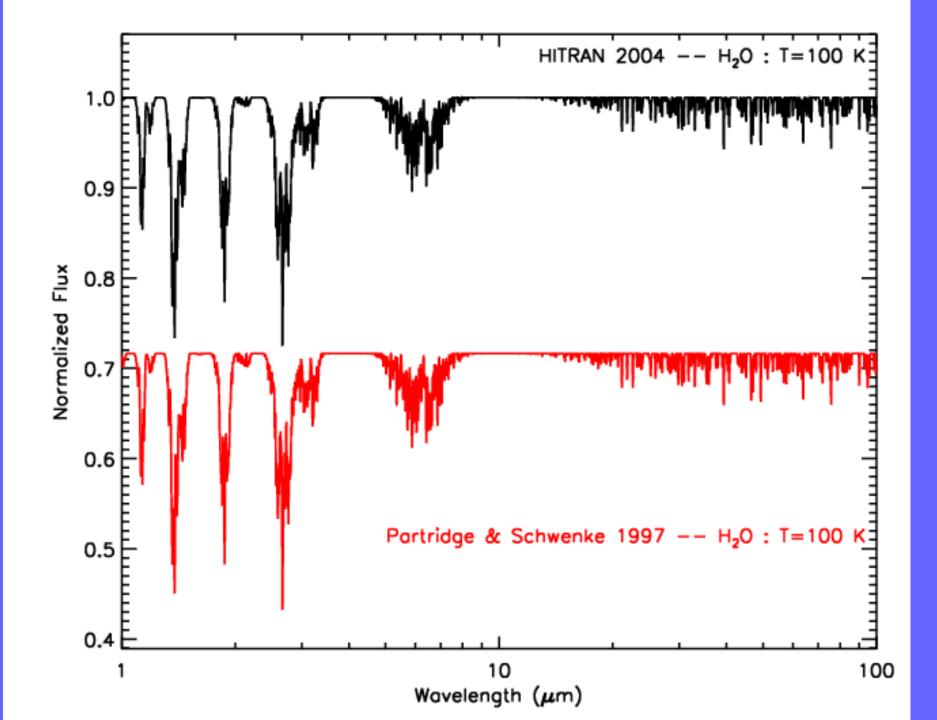
On request

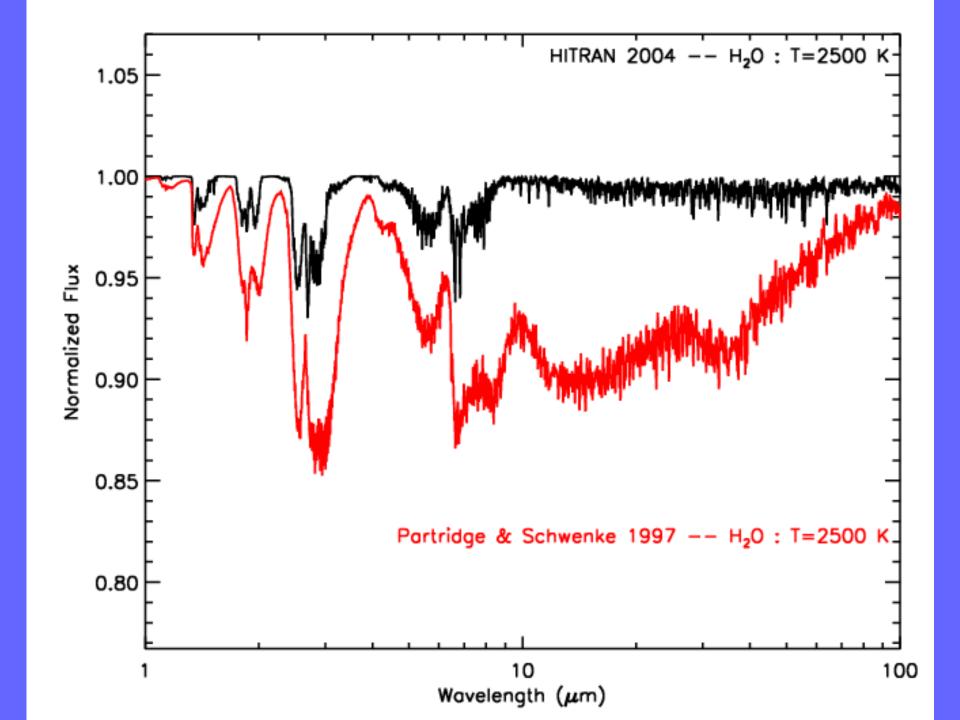
Robust statistics, fast

**Future** 

# **Molecular species**

- 143 molecules detected in space 212 including isotopologues most in sub-mm
  - 42 molecules studied in earth atmosphere line lists available at 296 K (HITRAN) often infrared
    - molecules with good line lists
       H<sub>2</sub>O, TiO, SiO, CO, CO<sub>2</sub>
       Large frequency range
       Many levels (electronic, vibrational, rotational)
       High temperatures

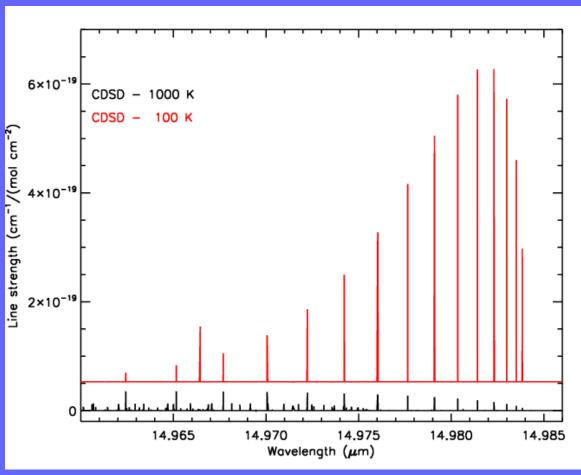




## **Model calculations I**

Line list: Frequencies of transition Intrinsic line strength at  $T_{ref}$ 

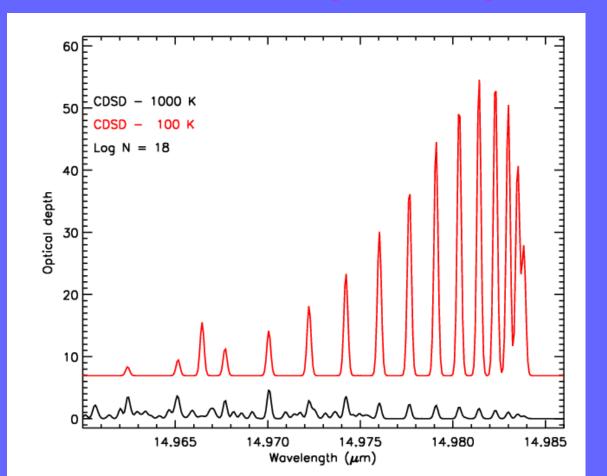
Calculate line strength at desired T



## **Model calculations II**

Multiply line strength by column density Convolve with intrinsic line profile

Calculate optical depth

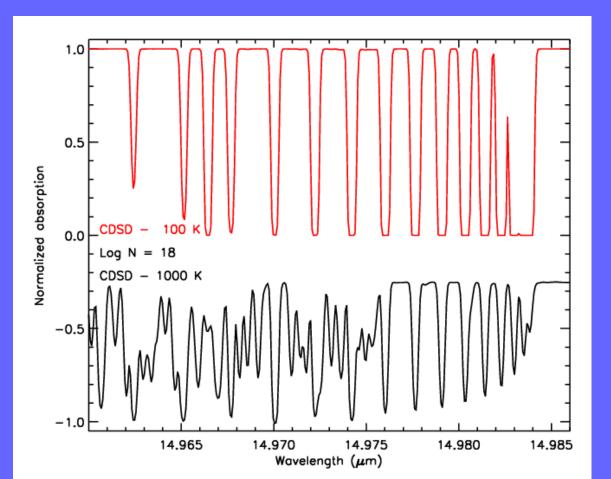


### **Model calculations III**

Assume LTE & slab

How does this absorb radiation?

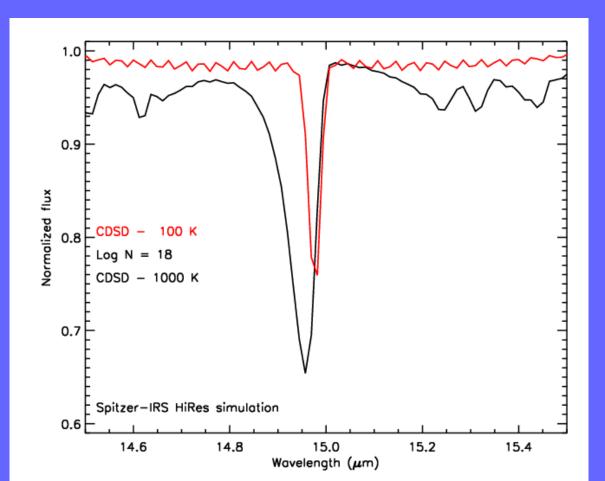
Radiative transfer



## **Model calculations IV**

#### **Convolve with instrument spectral response**

#### Instrument simulation



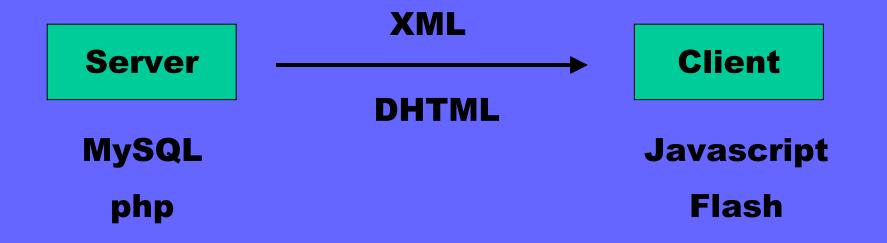
#### **Current contents of database**

- **42** molecules (95 isotopologues)
- 25 different temperatures (100 K 2500 K)
  - column densities (10<sup>16</sup> 10<sup>24</sup> cm<sup>-2</sup>)
  - **5** Instrument presets

291,213 model spectra (as of today)

Model spectra added by user requests !

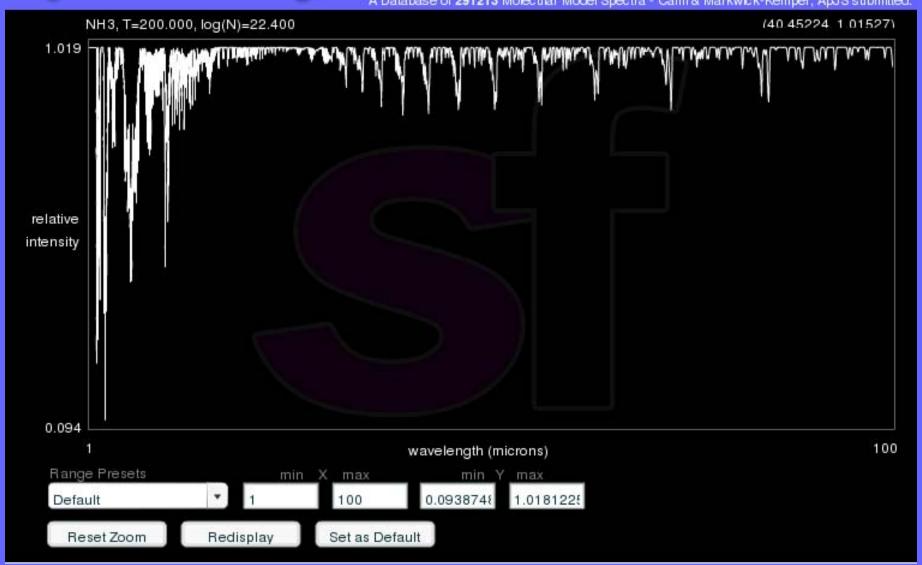
#### The web interface - architecture



http://www.spectrafactory.net/

# **spect**rafactory

A Database of 291213 Molecular Model Spectra - Cami & Markwick-Kemper, ApJS submitted.



| browsespectra          |   |
|------------------------|---|
| Molecule H2O           | Temperature 500 K to 500 K                                    |
| Isotopologue H2(16)O 💌 | log Column Density 18 cm <sup>-2</sup> to 18 cm <sup>-2</sup> |
| Line List All          |   |
| Instrument All         | Search  |

#### searchresults

select count(\*) from prv2 p where p.ml\_id=1 and p.is\_id=1 and temperature >= 500 and temperature <= 500 and log\_columndensity >= 18 and log\_columndensity <= 18

Found 15 spectra, showing 1 - 15

|    | Molecule         | Isotopologue                   | Line List | T/K | log N / cm <sup>-2</sup> | v / km s <sup>-1</sup> | Resolution | Oversample | Instrument           |
|----|------------------|--------------------------------|-----------|-----|--------------------------|------------------------|------------|------------|----------------------|
| 1  | H <sub>2</sub> O | н <sub>2</sub> <sup>16</sup> 0 | Schwenke  | 500 | 18                       | 3                      | 2000       | 2          | IRTF SpeX            |
| 2  | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | Schwenke  | 500 | 18                       | 3                      | 600        | 2          | Spitzer-IRS HiRes1   |
| 3  | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | HITEMP    | 500 | 18                       | 3                      | 120        | 2          | Spitzer-IRS LowRes2  |
| 4  | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | HITRAN04  | 500 | 18                       | 3                      | 120        | 2          | Spitzer-IRS LowRes2  |
| 5  | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | Schwenke  | 500 | 18                       | 3                      | 300        | 4          | ISO-SWS AOT1, speed1 |
| 6  | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | Schwenke  | 500 | 18                       | 3                      | 120        | 2          | Spitzer-IRS LowRes2  |
| 7  | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | HITRAN04  | 500 | 18                       | 3                      | 2000       | 2          | IRTF SpeX            |
| 8  | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | HITEMP    | 500 | 18                       | 3                      | 600        | 2          | Spitzer-IRS HiRes1   |
| 9  | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | HITRAN04  | 500 | 18                       | 3                      | 600        | 2          | Spitzer-IRS HiRes1   |
| 10 | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | HITEMP    | 500 | 18                       | 3                      | 90         | 2          | Spitzer-IRS LowRes1  |
| 11 | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | HITEMP    | 500 | 18                       | 3                      | 300        | 4          | ISO-SWS AOT1, speed1 |
| 12 | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | Schwenke  | 500 | 18                       | 3                      | 90         | 2          | Spitzer-IRS LowRes1  |
| 13 | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | HITRAN04  | 500 | 18                       | 3                      | 300        | 4          | ISO-SWS AOT1, speed1 |
| 14 | H <sub>2</sub> O | н <sub>2</sub> <sup>16</sup> 0 | HITRAN04  | 500 | 18                       | 3                      | 90         | 2          | Spitzer-IRS LowRes1  |
| 15 | H <sub>2</sub> O | H <sub>2</sub> <sup>16</sup> O | HITEMP    | 500 | 18                       | 3                      | 2000       | 2          | IRTF SpeX            |
|    |                  |                                |           |     |                          |                        |            |            |                      |

Found 15 spectra, showing 1 - 15

# **Needs & prospects**

More / better line lists

Inaccurate line lists better than no line list
Theoretical calculations and/or laboratory work
Availability is only limit to our database!

**Funding** 

This project is 0% funded

Observatories

Spitzer (now), soon SOFIA, Herschel, ALMA, ...

143 molecules detected in space; many more will be detected in the next years